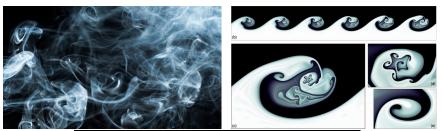
Turbulence Kolmogorov Phenomenology



Alexandros ALEXAKIS alexakis@phys.ens.fr Dep. Physique ENS Ulm

Richardson's Poem

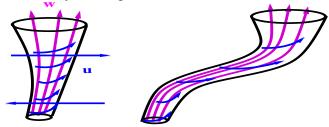


Big whorls have little whorls
Which feed on their velocity
And little whorls
have lesser whorls,
And so on to viscosity.

[Concerning atmospheric turbulence.]
Lewis Fry Richardson

Vortex stretching

Consider an eddy of velocity u_ℓ and lengthscale ℓ being sheared by an eddy of velocity U_L lengthscale L



The rate energy moves to smaller scales is:

$$\frac{dE_{\ell}}{dt} \propto \frac{U_L}{L} u_{\ell}^2$$

Vortex stretching



The rate energy moves to smaller scales is:

$$\frac{dE_{\ell}}{dt} \propto \frac{U_L}{L} u_{\ell}^2$$

Assuming

- (1) the flux of energy across scales is constant and equal to ϵ
- 2 the most effective interactions are among similar size eddies

$$\epsilon \propto \frac{u_\ell^3}{\ell}$$
 or $u_\ell \propto \epsilon^{1/3} \ell^{1/3}$



Fourier Space (Finite box)

$$\mathbf{u}(\mathbf{x},t) = \sum \tilde{\mathbf{u}}_{\mathbf{k}}(t)e^{i\mathbf{k}\cdot\mathbf{x}}, \qquad \tilde{\mathbf{u}}_{\mathbf{k}}(t) = \left\langle \mathbf{u}e^{-i\mathbf{k}\cdot\mathbf{x}} \right\rangle$$

Energy Spectrum:

$$E(k) = \frac{1}{2\delta k} \sum_{k \le |\mathbf{k}| < k + \delta k} |\tilde{\mathbf{u}}_{\mathbf{k}}|^2 \sim \frac{\text{Energy}}{\text{per unit wavenumber}}$$
$$k \propto 1/\ell, \qquad E(k)k \propto u_\ell^2$$

$$E(k) \propto \epsilon^{2/3} k^{-5/3}$$

Kolmogorov's Spectrum!



Kolmogorov scale

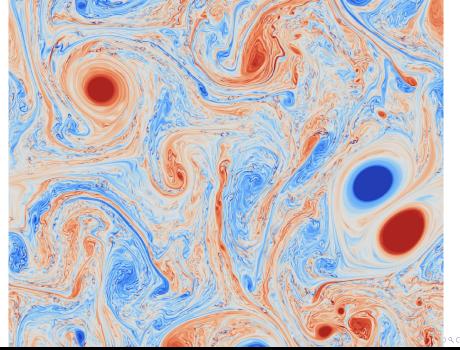
$$E(k) \propto \epsilon^{2/3} k^{-5/3}$$

Viscosity will become important when

$$\epsilon \propto \nu \frac{u_{\nu}^2}{\ell_{\nu}^2} \propto \nu \frac{(\epsilon^{1/3} \ell_{\nu}^{1/3})^2}{\ell_{\nu}^2} \propto \nu \frac{\epsilon^{2/3}}{\ell_{\nu}^{4/3}}$$

$$\boxed{\ell_{\nu} = \frac{\nu^{3/4}}{\epsilon^{1/4}}} \quad \text{or} \quad \boxed{k_{\nu} = \frac{\epsilon^{1/4}}{\nu^{3/4}}}$$





7/7