Banner appropriate to article type will appear here in typeset article

Bistability of the Large Scale Dynamics in Quasi-2D Turbulence

Xander M. de Wit^{1,2}, Adrian van Kan^{1,3} and Alexandros Alexakis¹†

¹Laboratoire de Physique de l'Ecole Normale Supérieure, ENS, Université PSL, CNRS, Sorbonne Université, Université Paris-Diderot, Sorbonne Paris Cité, Paris, France

(Received xx; revised xx; accepted xx)

In many geophysical and astrophysical flows, suppression of fluctuations along one direction of the flow drives a quasi-2D upscale flux of kinetic energy, leading to the formation of strong vortex condensates at the largest scales. Recent studies have shown that the transition towards this condensate state is hysteretic, giving rise to a limited bistable range in which both the condensate state as well as the regular 3D state can exist at the same parameter values. In this work, we use direct numerical simulations of thin layer flow to investigate whether this bistable range survives as the domain size and turbulence intensity are increased so that geophysically-relevant conditions are approached. By studying the time scales at which rare transitions occur from one state into the other, we find that the bistable range grows as the box size and/or Reynolds number *Re* are increased. We furthermore predict a crossover from a bimodal regime at low box size, low *Re* to a regime of pure hysteresis at high box size, high *Re*, in which any transition from one state to the other is prohibited at any finite time scale.

Key words:

1. Introduction

Ever since the seminal works of Batchelor (1969) and Kraichnan (1967), it has been known that in 2D turbulence, contrary to what is observed in 3D turbulence, kinetic energy cascades inversely, from the smaller scales at which it is injected to ever larger and larger scales. While the forward cascade that is observed in 3D turbulence is always arrested once it arrives at the scales at which viscosity becomes effective in dissipating the kinetic energy, such a stopping mechanism does not always exist at the large scales to saturate the inverse cascade. In that case, kinetic energy piles up at the largest available length scale of the flow system into what is referred to as a condensate. This condensate typically manifests as a strong vortex structure at the system size, also known as the Large-Scale Vortex, see figure 1.

Even in 3D flow systems, quasi-2D dynamics can be observed if fluctuations in one

²Fluids and Flows group, Department of Applied Physics and J. M. Burgers Centre for Fluid Dynamics, Eindhoven University of Technology, P.O. Box 513, 5600 MB Eindhoven, Netherlands

³Department of Physics, University of California, Berkeley, CA 94720, USA

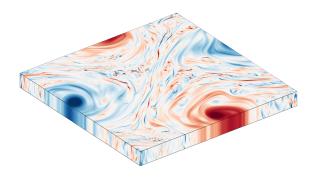


Figure 1: A vortex condensate in thin layer flow, visualised through a snapshot of vertical vorticity.

direction are strongly suppressed, allowing an inverse cascade to develop (Alexakis & Biferale 2018). In forced rotating turbulence (Biferale *et al.* 2016; Mininni *et al.* 2009; Smith *et al.* 1996; Smith & Waleffe 1999), rotating convection (Favier *et al.* 2014; Guervilly *et al.* 2014; Julien *et al.* 2012; Rubio *et al.* 2014) or rotating stratified turbulence (Pouquet & Marino 2013; Marino *et al.* 2013, 2014; van Kan & Alexakis 2020), such quasi-2D dynamics develops as a consequence of the Coriolis force that suppresses fluctuations along the axis of rotation according to the Taylor-Proudman theorem. Alternatively, such suppression could occur through for example magnetic forces (Alexakis 2011; Baker *et al.* 2018; Favier *et al.* 2010; Reddy *et al.* 2014) or plainly through geometric confinement as observed in thin layer flow (Celani *et al.* 2010; Benavides & Alexakis 2017; Musacchio & Boffetta 2017, 2019). This type of constrained dynamics is of eminent importance to many geo- and astrophysical flow settings, where (a combination of) the aforementioned mechanisms renders the flow quasi-2D. Examples can be found in our oceans (King *et al.* 2015; Scott & Wang 2005), in the atmosphere (Byrne & Zhang 2013; Nastrom *et al.* 1984) and on gas giant planets such as Jupiter and Saturn (Heimpel & Aurnou 2007; Heimpel *et al.* 2016; Stellmach *et al.* 2016).

This work focuses on the transition towards the condensate state of such quasi-2D systems. Remarkably, in spite of the inherently widely different nature of the considered flow systems, recent studies revealed that all across forced rotating turbulence (Alexakis 2015; Seshasayanan & Alexakis 2018; Yokoyama & Takaoka 2017), thin layer turbulence (van Kan & Alexakis 2019) and even the natural system of rotating convection (Favier *et al.* 2019; de Wit *et al.* 2021), the transition into the condensate is discontinuous and shows hysteresis. This gives rise to a limited bistable range in which both the quasi-2D condensate state and the 3D flow state can exist at the same parameters.

Since it is now known that this bistability can also survive in natural forcing conditions (Favier et al. 2019; de Wit et al. 2021) and the condensate can also form, albeit at more extreme parameters, between realistic no-slip walls (Aguirre Guzmán et al. 2020), we aim to investigate whether the bistable range in the condensate transition could also survive under parameter conditions that are relevant to geo- and astrophysical flows. Motivated by the remarkable similarities in the condensate transition across the different flow systems, we focus on the conceptually and computationally most basic system of forced thin layer turbulence. Specifically, we are interested in the dependence on the system size and the strength of turbulent forcing, quantified through the injection-scale Reynolds number Re, in order to investigate whether the bistable range of the condensate transition shrinks or grows as system size and Re are increased. We focus on very moderate values of the system size and

Re in order to be able to gather computationally demanding statistics about the bistable range and the rare transitions into and out of the condensate state. We then extrapolate our findings to the limits of large system size and large *Re* that are relevant to geo- and astrophysics in order to ultimately predict whether this bistable range could be observed in real-world natural flow settings.

2. Numerical approach

In order to study thin layer turbulence, we consider the idealised case of forced incompressible 3D flow in a triply periodic box of dimensions $L \times L \times H$, where the vertical direction is thin $H \ll L$. The flow system is identical to the one described in van Kan & Alexakis (2019). We consider a Cartesian coordinate system (x, y, z) with unit vectors (e_x, e_y, e_z) , where the thin vertical direction is chosen along e_z . The flow u(x, t) is then governed by the incompressible forced Navier-Stokes equations

$$\frac{\partial \boldsymbol{u}}{\partial t} + (\boldsymbol{u} \cdot \nabla)\boldsymbol{u} = -\nabla P + \nu \nabla^2 \boldsymbol{u} + \boldsymbol{f}, \qquad (2.1a)$$

$$\nabla \cdot \boldsymbol{u} = 0, \tag{2.1b}$$

where P(x,t) denotes the pressure divided by the constant density and v represents the kinematic viscosity of the fluid. We consider a stochastic forcing f(x,t) that is vertically invariant $(\partial f/\partial z = 0)$ and acts exclusively in the (e_x, e_y) plane, i.e. in the 2D2C (2-dimensional, 2-component) manifold. Furthermore, the forcing is divergence-free and acts only sharply on wavenumber $\dagger k_f \equiv 2\pi/\ell$, where its random phase is white noise (delta-correlated) in time. This results in a fixed mean injection rate $\langle u \cdot f \rangle = \epsilon$ that is solely prescribed by the forcing amplitude (Novikov 1965). Here, $\langle \cdot \rangle$ is used to represent the ensemble average. The choice of forcing is motivated by simplicity and comparability with previous studies. In general, one may consider various 3D forcing functions (Poujol *et al.* 2020).

The input parameters of the thin layer flow system are combined to give three dimensionless numbers. We define an injection-scale Reynolds number $Re \equiv (\epsilon \ell^4)^{1/3}/\nu$, the ratio between the forcing scale and the domain height $Q \equiv \ell/H$ and the ratio between the forcing scale and the width of the domain $K \equiv \ell/L$. Finally, we define a forcing time scale $\tau_f \equiv (\ell^2/\epsilon)^{1/3}$ and energy scale $E_f \equiv (\epsilon \ell)^{2/3}$ that are used to non-dimensionalise the different temporal and energetic quantities reported in this work, respectively.

Equations (2.1a)-(2.1b) are solved numerically in the triply periodic domain using a pseuspectral code that is an adapted version of the Geophysical High-Order Suite for Turbulence (Ghost) as introduced by Mininni *et al.* (2011), employing 2/3-dealiasing. In order to investigate the dependence of the condensate transition and its bistable range on Re and the box size, we take the results in van Kan *et al.* (2019) as a starting point and extend them to smaller and larger Re and K. For each value of K, Re we vary the thinness of the fluid layer Q as the principal control parameter in close vicinity to the condensate transition. An overview of the full set of input parameters that are considered in this work is provided in table 1.

Resolutions $N_x \times N_y \times N_z$ are chosen such that we maintain the same ratio between the grid spacing $L/N_{x,y}$ and Kolmogorov length $\eta = (v^3/\epsilon)^{1/4}$ as used in van Kan *et al.* (2019) of $L/N_{x,y} \approx 3.2\eta$ in the horizontal directions and we resolve finer than that in the vertical direction. For the vertical, we ensure that we keep 16 grid cells in order to maintain sufficient degrees of freedom in the thin direction.

and varying Ke (right).								
1/K	Re	Q	$N_x \times N_y \times N_z$		1/K	Re	Q	$N_x \times N_y \times N_z$
6	192	[1.53:1.70]	96 × 96 × 16		8	76	[1.00 : 1.13]	$64 \times 64 \times 16$
7	192	[1.53:1.69]	$112 \times 112 \times 16$		8	131	[1.29:1.48]	$96 \times 96 \times 16$
8	192	[1.44:1.81]	$128 \times 128 \times 16$		8	192	[1.44:1.81]	$128 \times 128 \times 16$
9	192	[1.44 : 1.73]	$144 \times 144 \times 16$		8	329	[1.74:2.05]	$192 \times 192 \times 16$

Table 1: The different series of input parameters used in this work for varying box size (left) and varying *Re* (right).

For each unique set of input parameters (Q, K, Re), at least 40 independent runs are carried out by using a different random seed for the stochastic forcing in order to obtain statistics about the rare transitions from one state into the other. The combination of this need for a multitude of independent runs with the fact that the large-scale dynamics is very slow compared to the background smaller-scale 3D dynamics renders only the rather moderate parameters that are considered here computationally accessible. The simulations in this work comprise more than ~ 1.0 million CPU hours.

The main diagnostic quantity that we use here to probe the strength of the condensate is the 2D large-scale energy E_{ls} , defined from the Fourier components $\hat{u}(k)$ of the flow as

$$E_{ls} = \frac{1}{2} \sum_{\substack{\boldsymbol{k} \\ |\boldsymbol{k}| \leq k_{\text{max}} \\ k_{-} = 0}} \left(|\hat{\boldsymbol{u}}(\boldsymbol{k}) \cdot \boldsymbol{e}_{x}|^{2} + |\hat{\boldsymbol{u}}(\boldsymbol{k}) \cdot \boldsymbol{e}_{y}|^{2} \right), \tag{2.2}$$

where the cut-off wavenumber $k_{\text{max}} = \sqrt{2}(2\pi/L)$.

3. Bistability, bimodality and rare transitions

Bistability and bimodality are often met in dynamical systems. Here we refer to bistability as the presence of two independent stable attractors, coexisting for the same value of parameters, whilst by bimodality we refer to the case that two attractors are linked by some trajectories that when followed the system jumps from one attractor to the other and vice-versa. In the former case, a hysteresis loop can exist when one of the parameters of the system is continuously varied. In an inherently fluctuating dynamical system like the one at hand however, the precise extent of the bistable range in the hysteresis loop can be ambiguous as rare sudden transitions from one hysteretic branch into the other may exist at very long time scales near both ends of the hysteresis loop. This raises the question how we can unambiguously define the precise extent of the bistable range in the hysteretic transition.

Earlier works that studied the bistable range in the quasi-2D condensate transition would typically simulate up to a certain time scale that is constrained by computational limits and call a state stable if no further transitions are observed (Favier *et al.* 2019; van Kan & Alexakis 2019; de Wit *et al.* 2021; Yokoyama & Takaoka 2017). However, in the view of the rare transitions, this amounts to an in principle arbitrary cut-off at a certain time scale, neglecting any possible transitions occurring at larger time scales. We refer to this as *finite-time hysteresis*.

However, in the case of *pure hysteresis*, in strict sense, all transitions from one state into the other are prohibited at any finite time scale, such that the system is absolutely bistable. This would require the time scales of the rare transitions to diverge at a certain asymptote. The existence of such asymptotes is the principal assumption in this work. As we will show

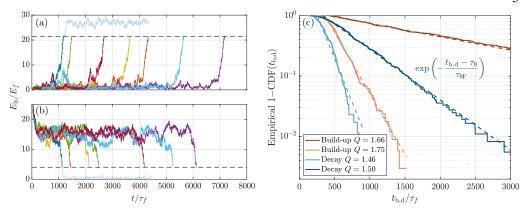


Figure 2: Left: Examples of time series of different realisations for build-up (a) and decay (b) (at Q=1.67 (a) and Q=1.49 (b) and 1/K=9, Re=192). Horizontal dashed lines represent the threshold energy at which the build-up time t_{D} or decay time t_{d} is defined and the simulation is terminated. First blue run is continued for demonstration purposes. Right: Examples of distributions of $t_{D,d}$, shown through the empirical cumulative density function (at 1/K=8 and Re=192). Dashed lines represent fits of the exponential distribution.

in section 4, we can define such asymptotes based on the scaling of the time scales of the rare transitions that we observe, allowing us to study how these asymptotes shift as we vary the box size and *Re* in order to get a completely time scale independent and unambiguous method for comparing our results at these different parameters.

To analyse the rare transitions between both states, we separately consider build-up events from the 3D state into the condensate state and decay events vice-versa. The build-up events are studied by initialising the simulations with a tiny perturbation onto a state of no flow and continuing until the condensate state is reached. For the decay events, we initialise the simulation with a snapshot from a condensate state at higher Q and we continue the run until the condensate has decayed and the 3D state is obtained.

Figures 2a-b show examples of different realisations of such rare transitions for one choice of parameters. The works of van Kan et al. (2019) and de Wit et al. (2021) have revealed that the waiting time that is spent until these sudden transitions from one state into the other commence is exponentially distributed, signifying that the transition process is memoryless. The typical mean waiting time τ_W can be obtained by defining representative thresholds in the large-scale energy and analysing the distribution of times $t_{\rm b,d}$ after which these thresholds are crossed. Examples of the obtained empirical cumulative distributions are depicted in figure 2c, showing that it closely follows the aforementioned exponential distribution. We can then obtain τ_W by fitting the empirical cumulative distribution function CDF with a (shifted) exponential as

$$CDF(t_{b,d}) = 1 - \exp\left(-\frac{t_{b,d} - \tau_0}{\tau_W}\right). \tag{3.1}$$

This process can be repeated to obtain the waiting time scales τ_W for build-up events and decay events at different values of Q, varying it across the full extent of the hysteretic transition. This results in a series of typical waiting times τ_W as a function of Q, which can in turn be repeated for different box sizes and Re.

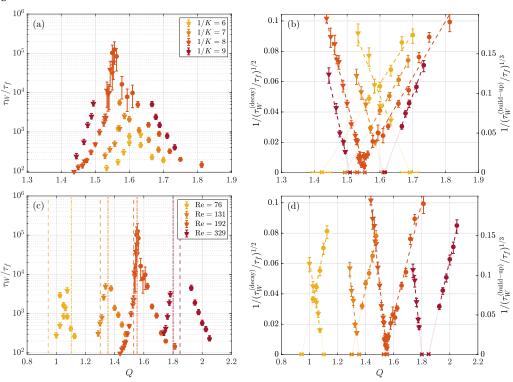


Figure 3: Waiting times τ_W (a,c) for build-up (circles) and decay (triangles) events and their power law transformation (b,d) for varying box size 1/K at Re=192 (a,b) and varying Re at 1/K=8 (c,d). Crosses on the horizontal axis in (b,d) denote the estimates for the asymptotes Q_0 . These asymptotes are also depicted in (c) by the vertical lines for build-up (dashed) and decay (dashed-dotted), but are omitted in (a) for readability.

4. Time-scale statistics

The results for the series of transition time scales for varying box size and varying Re are provided in the left panels of figure 3. As the transition is approached τ_W increases faster than exponentially (see van Kan et al. 2019). This implies that either τ_W increases in a non-diverging superexponential fashion (e.g. $\tau_W \propto \exp[\exp(Q)]$, as is typical for certain transitions controlled by extreme events (Goldenfeld et al. 2010; Nemoto & Alexakis 2018, 2021; Gomé et al. 2021)), or that it diverges at some critical value Q_0 . To determine which of the two holds for the present system is beyond the scope of this work. We will thus assume that the latter case applies, although one could alternatively interpret Q_0 as the value at which super-exponential behaviour starts in the former case. To determine Q_0 we fit the transition time to a power-law divergence

$$\tau_W \propto \frac{1}{|Q - Q_0|^p}.\tag{4.1}$$

By plotting $1/\tau_W^{1/p}$ as a function of Q, we can then obtain Q_0 from a linear fit to the data (right panels of figure 3). Empirically, we find that $p^{\text{(build-up)}} = 3$ and $p^{\text{(decay)}} = 2$ result in a satisfactory linearisation of our data. These asymptotes then predict the location in Q where the transition time becomes infinite, such that we can say that beyond the asymptote, the transition can not occur at any finite time scale.

Comparing the results at different box sizes and different Re, we find first of all that the

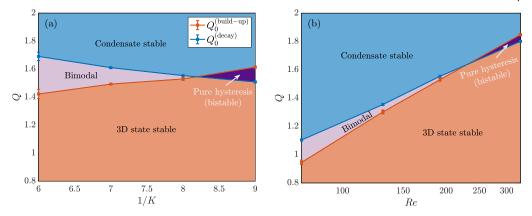


Figure 4: Phase diagrams of the condensate for varying box size at Re = 192 (a) and varying Re at 1/K = 8 (b). Red and blue lines denote the asymptotes of waiting times for build-up and decay events, respectively.

transition is observed in a similar range of Q for the varying box sizes, while it clearly shifts as Re is varied, in agreement with the observations in van Kan & Alexakis (2019). More importantly, we observe that the branches of build-up and decay times move further apart as Re and box-size are increased. The branches cross for the runs with $Re \leqslant 192$ or for $K \leqslant 8$, such that at small box size and/or small Re, a bimodal range of Q exists for $Q_0^{\text{(build-up)}} < Q < Q_0^{\text{(decay)}}$ where the build-up and decay time scales are simultaneously finite (and in fact computationally accessible). Hence in this range, the flow continually transitions back and forth between the 3D state and the condensate state.

Conversely, for the largest box size and largest Re that we consider, the branches of the build-up and decay branches never cross as the asymptotes reside on opposite ends. This indicates a profoundly different regime, where the decay times have diverged before the build-up times become finite, such that in the range $Q_0^{(\text{decay})} < Q < Q_0^{(\text{build-up})}$, both states are absolutely stable as no transition from one state into the other can occur in any finite time. This corresponds to a regime of *pure hysteresis* in which the system is (absolutely) bistable. Indeed, it is this range that is arguably the most unambiguous time-scale independent definition of the bistable range of the system. These results are summarised in figure 4. It is thus evident from our results that the bistable range grows as the box size and/or Re are independently increased.

We argue that the strengthening of the bistability for large box sizes and Re is intuitive from the increase of the condensate energy level (compared to the 3D-state energy) as Re, 1/K is increased, making it harder to jump from one state into the other. Our results thus indicate that bistability is not a finite size, finite Re effect and such states can be found in the geophysical limit where both Re and domain size are large.

5. Conclusions and outlook

In this work, we have demonstrated that the bistability observed in the transition to the quasi-2D condensate state can survive under the geophysically relevant conditions of increasing box size and increasing Re. By studying the time scales at which rare transitions from one state into the other occur we quantified the precise extent of the bistable range. Fitting the mean time scales of these transitions with a diverging power law, we measured the locations of the asymptotes beyond which the transition is prohibited at any finite time scale. Since these asymtotes show a crossover as we vary the box size and/or Re, this predicts a profound regime

change from a bimodal regime at small box size and/or *Re* to a regime of pure hysteresis at large box size and/or *Re*, as summarised in figure 4. Since we find that the branches of time scales of build-up transitions into the condensate and decay transition out of the condensate at both ends of the bistable range only separate further and further as the box size and/or *Re* is increased, we conclude that this bistability is not a finite size or finite *Re* effect, but that the bistable range grows as we progress towards the geophysical limit.

We remark that the method proposed here for quantifying the precise extent of the bistable range in a hysteretic transition using the time scales of rare transitions is entirely general and a similar procedure can be followed in the context of any other fluctuating hysteretic dynamical system within or beyond fluid dynamics. However, we must note that the motivation of (4.1) is ad hoc here. Although the agreement with our data as shown in 3 is satisfactory, a more fundamental physical motivation, supported by a larger dynamic range of parameters, would be needed to rigorously prove the validity of (4.1) as well as our choice of exponents. Indeed, although the waiting time increases faster than exponentially, the existence of an asymptote for τ_W in the first place is ultimately an assumption in itself and the possibility of the relation being any other superexponential relation without divergence can in principle not be ruled out by numerical simulations alone. Nonetheless, while the existence of pure hysteresis certainly constrains the underlying physical mechanism of the transition, one may argue that it does not hold immediate implications in geophysical practice whether the time scales of transitions are strictly infinite or merely beyond any practical finite time scale. Moreover, we argue that the satisfactory agreement of (4.1) with our data in itself does convincingly prove our central result: that the bistable range is not an effect of finite box size and/or finite Re, but that it grows as we progress towards the geophysical conditions. Indeed, this holds either in the strict terms of absolute bistability, or in the terms of exceeding a certain finite superexponential time scale.

While recent investigations have started to unveil different aspects of this peculiar type of transition between turbulent flow states, much of the underlying physical mechanism still remains poorly understood. In particular, which specific physical events trigger the flow to commence the transition, for example either a series of vortex merging events, or rare fluctuations directly at the largest scale, remains an open question. Answering such questions would contribute greatly to our understanding of this flow phenomenon and our work may act as a numerical inspiration as well as a quantitative benchmark to such theoretical studies. Specifically, understanding the physical mechanism behind the transition may motivate the theoretical validity of relation (4.1), which we have motivated only empirically here.

In order to push the analysis in this work further towards the geophysical conditions of yet larger box size and *Re*, direct numerical simulations quickly become computationally unfeasible, as both spatial resolution requirements as well as time scale requirements increase rapidly. A promising solution may lie in the application of rare event algorithms (Cérou & Guyader 2007; Lestang *et al.* 2018). Such algorithms are more efficient in probing rare transitions and have been successfully applied in various other flow contexts (Gomé *et al.* 2021; Bouchet *et al.* 2019; Rolland 2018). By progressing further towards the extreme geophysical conditions, we may for example investigate whether the growth of the bistable range of the condensate transition saturates at some point, which can not be studied from the moderate parameters considered in our work. Finally, we remark that it also seems attractive now to study the condensate transition and its bistable behaviour from experiments in which more extreme parameters may be more easily accessible, or perhaps even from observations in real-world geophysical or astrophysical flows.

Acknowledgments. The authors thank H.J.H. Clercx for useful discussions and comments.

Funding. This work was granted access to the HPC resources of MesoPSL financed by the Region Ile de

France and the project Equip@Meso (reference ANR-10-EQPX-29-01) of the programme Investissements d'Avenir supervised by the Agence Nationale pour la Recherche and the HPC resources of GENCI-TGCC & GENCI-CINES (Project No. A0090506421, A0110506421). This work has also been supported by the Agence nationale de la recherche (ANR DYSTURB project No. ANR-17-CE30-0004). AvK acknowledges support by Studienstiftung des deutschen Volkes and the National Science Foundation (grant DMS-2009563).

Declaration of Interests. The authors report no conflict of interest.

REFERENCES

- AGUIRRE GUZMÁN, A. J., MADONIA, M., CHENG, J. S., OSTILLA-MÓNICO, R., CLERCX, H. J. H. & KUNNEN, R. P. J. 2020 Competition between Ekman Plumes and Vortex Condensates in Rapidly Rotating Thermal Convection. *Phys. Rev. Lett.* **125** (21), 214501.
- ALEXAKIS, A. 2011 Two-dimensional behavior of three-dimensional magnetohydrodynamic flow with a strong guiding field. *Phys. Rev. E* **84** (5), 056330.
- ALEXAKIS, A. 2015 Rotating Taylor-Green flow. J. Fluid Mech. 769, 46–78.
- ALEXAKIS, A. & BIFERALE, L. 2018 Cascades and transitions in turbulent flows. Phys. Rep. 767-769, 1-101.
- BAKER, N. T., POTHÉRAT, A., DAVOUST, L. & DEBRAY, F. 2018 Inverse and Direct Energy Cascades in Three-Dimensional Magnetohydrodynamic Turbulence at Low Magnetic Reynolds Number. *Phys. Rev. Lett.* **120** (22), 224502.
- BATCHELOR, G. K. 1969 Computation of the energy spectrum in homogeneous two-dimensional turbulence. *Phys. Fluids* **12** (12), II–233.
- Benavides, S. J. & Alexakis, A. 2017 Critical transitions in thin layer turbulence. *J. Fluid Mech.* **822**, 364–385.
- BIFERALE, L., BONACCORSO, F., MAZZITELLI, I. M., VAN HINSBERG, M. A. T., LANOTTE, A. S., MUSACCHIO, S., PERLEKAR, P. & TOSCHI, F. 2016 Coherent structures and extreme events in rotating multiphase turbulent flows. *Phys. Rev. X* **6** (4), 1–14.
- BOUCHET, F., ROLLAND, J. & SIMONNET, E. 2019 Rare Event Algorithm Links Transitions in Turbulent Flows with Activated Nucleations. *Phys. Rev. Lett.* **122** (7), 074502.
- Byrne, D. & Zhang, J. A. 2013 Height-dependent transition from 3-D to 2-D turbulence in the hurricane boundary layer. *Geophys. Res. Lett.* **40** (7), 1439–1442.
- Celani, A., Musacchio, S. & Vincenzi, D. 2010 Turbulence in More than Two and Less than Three Dimensions. *Phys. Rev. Lett.* **104** (18), 184506.
- CÉROU, F. & GUYADER, A. 2007 Adaptive Multilevel Splitting for Rare Event Analysis. *Stoch. Anal. Appl.* **25** (2), 417–443.
- Favier, B., Godeferd, F. S., Cambon, C. & Delache, A. 2010 On the two-dimensionalization of quasistatic magnetohydrodynamic turbulence. *Phys. Fluids* 22 (7), 075104.
- Favier, B., Guervilly, C. & Knobloch, E. 2019 Subcritical turbulent condensate in rapidly rotating Rayleigh-Bénard convection. *J. Fluid Mech.* **864**, R1.
- FAVIER, B., SILVERS, L. J. & PROCTOR, M. R. E. 2014 Inverse cascade and symmetry breaking in rapidly rotating Boussinesq convection. *Phys. Fluids* **26** (9), 096605.
- GOLDENFELD, N., GUTTENBERG, N. & GIOIA, G. 2010 Extreme fluctuations and the finite lifetime of the turbulent state. *Phys. Rev. E* **81**, 035304.
- Gomé, S., Tuckerman, L. S. & Barkley, D. 2021 Extreme events in transitional turbulence. (pre-publication), arXiv: 2109.01476.
- Guervilly, C., Hughes, D. W. & Jones, C. A. 2014 Large-scale vortices in rapidly rotating Rayleigh-Bénard convection. *J. Fluid Mech.* **758**, 407–435.
- HEIMPEL, M. & AURNOU, J. 2007 Turbulent convection in rapidly rotating spherical shells: A model for equatorial and high latitude jets on Jupiter and Saturn. *Icarus* **187** (2), 540–557.
- HEIMPEL, M., GASTINE, T. & WICHT, J. 2016 Simulation of deep-seated zonal jets and shallow vortices in gas giant atmospheres. *Nat. Geosci.* **9** (1), 19–23.
- Julien, K., Rubio, A. M., Grooms, I. & Knobloch, E. 2012 Statistical and physical balances in low Rossby number Rayleigh–Bénard convection. *Geophys. Astrophys. Fluid Dyn.* **106** (4-5), 392–428.
- VAN KAN, A. & ALEXAKIS, A. 2019 Condensates in thin-layer turbulence. J. Fluid Mech. 864, 490-518.
- VAN KAN, A. & ALEXAKIS, A. 2020 Critical transition in fast-rotating turbulence within highly elongated domains. *J. Fluid Mech.* **899**, A33.

- VAN KAN, A., Nемото, T. & Alexakis, A. 2019 Rare transitions to thin-layer turbulent condensates. *J. Fluid Mech.* 878, 356–369.
- King, G. P., Vogelzang, J. & Stoffelen, A. 2015 Upscale and downscale energy transfer over the tropical Pacific revealed by scatterometer winds. *J. Geophys. Res. Oceans* **120** (1), 346–361.
- Kraichnan, R. H. 1967 Inertial ranges in two-dimensional turbulence. *Phys. Fluids* 10 (7), 1417–1423.
- Lestang, T., Ragone, F., Bréhier, C.-E., Herbert, C. & Bouchet, F. 2018 Computing return times or return periods with rare event algorithms. *J. Stat. Mech.: Theory Exp.* **4** (4), 043213.
- MARINO, R., MININNI, P. D., ROSENBERG, D. & POUQUET, A. 2013 Inverse cascades in rotating stratified turbulence: Fast growth of large scales. *EPL* **102** (4), 44006.
- MARINO, R., MININNI, P. D., ROSENBERG, D. L. & POUQUET, A. 2014 Large-scale anisotropy in stably stratified rotating flows. *Phys. Rev. E* **90** (2), 023018.
- MININNI, P. D., ALEXAKIS, A. & POUQUET, A. 2009 Scale interactions and scaling laws in rotating flows at moderate Rossby numbers and large Reynolds numbers. *Phys. Fluids* **21** (1), 015108.
- MININNI, P. D., ROSENBERG, D., REDDY, R. & POUQUET, A. 2011 A hybrid MPI–OpenMP scheme for scalable parallel pseudospectral computations for fluid turbulence. *Parallel Comput.* **37** (6-7), 316–326.
- Musacchio, S. & Boffetta, G. 2017 Split energy cascade in turbulent thin fluid layers. *Phys. Fluids* **29** (11), 111106.
- Musacchio, S. & Boffetta, G. 2019 Condensate in quasi-two-dimensional turbulence. *Phys. Rev. Fluids* **4** (2), 022602(R).
- Nastrom, G. D., Gage, K. S. & Jasperson, W. H. 1984 Kinetic energy spectrum of large-and mesoscale atmospheric processes. *Nature* **310** (5972), 36–38.
- Nемото, Т. & Alexakis, A. 2018 Method to measure efficiently rare fluctuations of turbulence intensity for turbulent-laminar transitions in pipe flows. *Phys. Rev. E* 97, 022207.
- Nемото, Т. & Alexakis, A. 2021 Do extreme events trigger turbulence decay? a numerical study of turbulence decay time in pipe flows. *J. Fluid Mech.* **912**, A38.
- Novikov, E. A. 1965 Functionals and the random-force method in turbulence theory. *Sov. Phys. JETP* **20** (5), 1290–1294.
- Poujol, B., van Kan, A. & Alexakis, A. 2020 Role of the forcing dimensionality in thin-layer turbulent energy cascades. *Phys. Rev. Fluids* **5** (6), 064610.
- Pouquet, A. & Marino, R. 2013 Geophysical turbulence and the duality of the energy flow across scales. *Phys. Rev. Lett.* **111** (23), 234501.
- REDDY, K. S., KUMAR, R. & VERMA, M. K. 2014 Anisotropic energy transfers in quasi-static magnetohydrodynamic turbulence. *Phys. Plasmas* 21 (10), 102310.
- Rolland, J. 2018 Extremely rare collapse and build-up of turbulence in stochastic models of transitional wall flows. *Phys. Rev. E* **97** (2), 023109.
- Rubio, A. M., Julien, K., Knobloch, E. & Weiss, J. B. 2014 Upscale Energy Transfer in Three-Dimensional Rapidly Rotating Turbulent Convection. *Phys. Rev. Lett.* **112** (14), 144501.
- Scott, R. B. & Wang, F. 2005 Direct Evidence of an Oceanic Inverse Kinetic Energy Cascade from Satellite Altimetry. *J. Phys. Oceanogr.* **35** (9), 1650–1666.
- Seshasayanan, K. & Alexakis, A. 2018 Condensates in rotating turbulent flows. *J. Fluid Mech.* **841**, 434–462.
- Smith, L. M., Chasnov, J. R. & Waleffe, F. 1996 Crossover from Two- to Three-Dimensional Turbulence. *Phys. Rev. Lett.* **77** (12), 2467–2470.
- Smith, L. M. & Waleffe, F. 1999 Transfer of energy to two-dimensional large scales in forced, rotating three-dimensional turbulence. *Phys. Fluids* **11** (6), 1608–1622.
- STELLMACH, S., VERHOEVEN, J., LISCHPER, M. & HANSEN, U. 2016 Towards a Better Understanding of Rotating Turbulent Convection in Geo- and Astrophysical Systems. In *NIC Symposium 2016*. Jülich.
- DE WIT, X. M., AGUIRRE GUZMÁN, A. J., CLERCX, H. J. H. & KUNNEN, R. P. J. 2021 Discontinuous Transitions Towards Vortex Condensates in Buoyancy-Driven Rotating Turbulence: Analogies with First-Order Phase Transitions. (*in review*), arXiv: 2106.01158.
- Yокоуама, N. & Такаока, M. 2017 Hysteretic transitions between quasi-two-dimensional flow and three-dimensional flow in forced rotating turbulence. *Phys. Rev. Fluids* **2** (9), 092602(R).